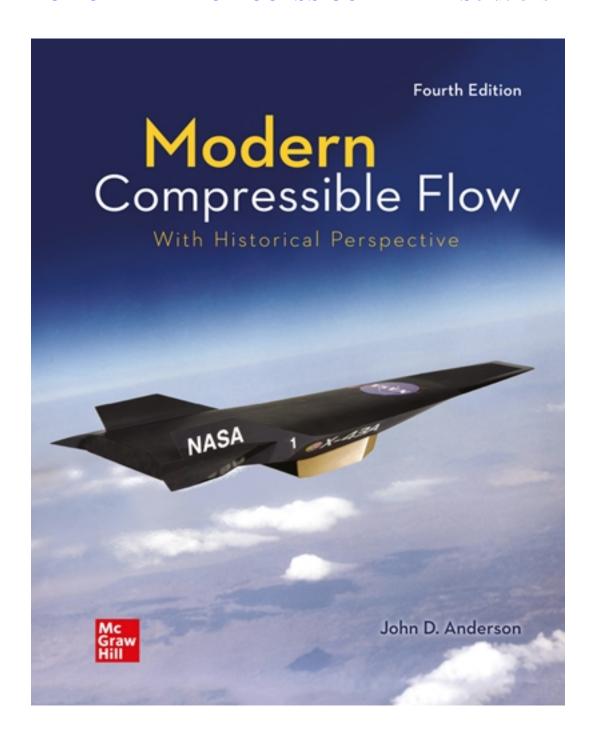
Solutions for Modern Compressible Flow 4th Edition by Anderson

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Solutions

CHAPTER 1

1.1
$$\rho = \frac{p}{RT} = \frac{(5.6)(2116)}{(1716)(850)} = \boxed{0.00812 \text{ slug/ft}^3}$$

$$v = \frac{1}{\rho} = 123 \text{ ft}^3/\text{slug}$$

1.2
$$\rho = \frac{p}{RT} = \frac{(10)(1.01 \text{ x } 10^5)}{(287)(320)} = \boxed{11.0 \text{ kg/m}^3}$$

$$n = \frac{p}{RT} = \frac{(10)(1.01 \times 10^5)}{(1.38 \times 10^{-23})(320)} = \boxed{2.87 \times 10^{26}/\text{m}^3}$$

$$\eta = \frac{pv}{RT} = \frac{p}{\rho RT} = \frac{(10)(1.01 \text{ x } 10^5)}{(11.0)(8314)(320)} = 0.0345 \frac{\text{kg} - \text{mole}}{\text{kg}}$$

1.3 From the definition of enthalpy,

$$h = e + p v = e + RT \tag{A1}$$

For a calorically perfect gas, this becomes

$$c_p T = c_v T + RT$$
, or $c_p - c_v = R$

For a thermally perfect gas, Eq. (A1) is first differentiated

$$dh = de + Rdt$$

or,

$$c_p dT = c_v dT = Rdt$$

or,

$$c_p - c_v = R$$

1.4
$$s_2 - s_1 = c_p \, \ell n \, \frac{T_2}{T_1} - R \, \ell n \, \frac{p_2}{p_1}$$

(a) $R = 1716 \text{ ft-lb/slug} \circ R$

$$c_p = \frac{\gamma R}{\gamma - 1} = \frac{(1.4)(1716)}{0.4} = 6006 \text{ ft-lb/slug}^{\circ}R$$

$$s_2 - s_1 = (6006) \ \ell n \ (1.687) - (1716) \ \ell n \ 4.5$$

$$s_2 - s_1 = 559.9 \text{ ft-lb/slug}^{\circ} R$$

(b) $R = 287 \text{ joule/kg}^{\circ}K$

$$c_p = \frac{\gamma R}{\gamma - 1} = \frac{(1.4)(287)}{0.4} = 1004.5 \text{ joule/kg}^{\circ} K$$

$$s_2 - s_1 = (1004.5) \ \ell n \ (1.687) - 287 \ \ell n \ (4.5)$$

$$s_2 - s_1 = 93.6 \text{ joule/kg}^{\circ}\text{K}$$

1.5
$$\frac{p_2}{p_1} = \left(\frac{T_2}{T_1}\right)^{\frac{\gamma}{\gamma - 1}}$$

$$p_2 = p_1 \left(\frac{T_2}{T_1}\right)^{\frac{\gamma}{\gamma - 1}} = 1800 (400/500)^{\frac{1.4}{0.4}}$$

$$p_2 = 824.3 \text{ lb/ft}_2$$

$$\rho_2 = \frac{p_2}{RT_2} = \frac{824.3}{(1716)(400)} = \boxed{0.0012 \text{ slug/ft}^3}$$

1.6 Volume of room = $(20)(15)(8) = 2400 \text{ ft}^3$

Standard sea level density = $0.002377 \text{ slug/ft}^3$

Mass of air =
$$(0.002377)(2400) = 5.70 \text{ slug}$$

Weight = Mass x acceleration of gravity = (5.7)(32.2) = 184 lb

1.7 (a)
$$dp = -\rho V dV$$

and
$$d\rho = \rho \tau dp$$
, or $dp = \frac{d\rho}{\rho \tau}$

Combining:

$$\frac{\mathrm{d}\rho}{\rho\tau} = -\rho \,\mathrm{VdV}$$

$$d\rho = -\tau p^2 V dV$$

$$\frac{\mathrm{d}\rho}{\rho} = -\tau \rho \,\mathrm{VdV}$$

$$\frac{\mathrm{d}\rho}{\rho} = -\tau \,\rho \,\mathrm{V}^2 \,\frac{\mathrm{d}\mathrm{V}}{\mathrm{V}}$$

(b)
$$\tau_s = \frac{1}{\gamma p} = \frac{1}{(1.4)(1.01 \text{ x } 10^5)} = 7.07 \text{ x } 10^{-6} \text{ m}^2/\text{N}$$

$$\frac{d\rho}{\rho} = \tau_s \rho V^2 \frac{dV}{V} = -(7.07 \times 10^{-6})(1.23)(10)^2(0.01)$$

$$\frac{d\rho}{\rho} = -8.7 \times 10^{-6}$$

(c) Here, $\frac{d\rho}{\rho}$ will be larger by the ratio $\left(\frac{1000}{10}\right)^2$.

$$\frac{d\rho}{\rho} = (-8.7 \times 10^{-6}) \left(\frac{1000}{10}\right)^2 = -8.7 \times 10^{-2}$$

<u>Comment</u>: By increasing the velocity of a factor of 100, the fractional change in density is increased by factor of 10⁴. This is just another indication of why high-speed flows must be treated as compressible.